BCS introduction

Getting access to full text astronomical articles.

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25th September 2023



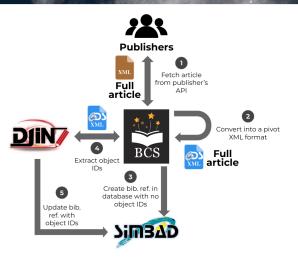
□ 1. BCS



Bibliographic **C**enter **S**upervisor (*CDS internal service*)

- 1. Get complete articles from astronomical journals
 - download articles (ideally in XML)
 - convert them into an internal XML format: *XCDS*
- 2. Give access to:
 - articles description
 - full text articles 🂵
 - tables (mainly, MRT)

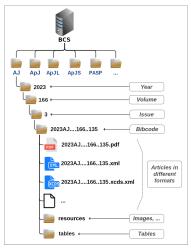
□ 2. How does it work?



(cf LISA IX: Poster, Video, Article)

□ 3. In practice

Article organization

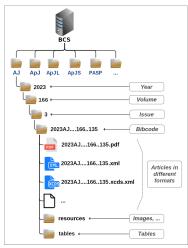


Ingestion management

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□ 3. In practice

Article organization



Ingestion management

Volume/Issue to import			
166_3 V Import in SIMBAD			Check authors Parfile saved
Volume: 166 Issue: 3	# articles: 55		
Fetched: 9/6/2023 4/29/06 PM		(on 559 authors)	
0	New authors Parfile 🥒	MaJ 🔒	
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Page: 81			
Nb pages: 20			
DOI: 10.3847/1538-3881/ace443			
ISSN: 1538-3881			
Copyright: © 2023. The Author(s). Published by the American	Astronomical Society.		
Date: 2023-090			
Authors: SALZER L.J. 0			
CARR D.J. O			
SIEDEN J. O			
BRUNKER S.W.			
HIRSCHAUER A.S. O			
Title: The Star Formation Across Cosmic Time (SFACT) S	urvey. I. Survey Description and Early R	esuits from a New Na	rrowband Emission-line Galaxy Survey.
Abstract: We introduce the Star Formation Arrows Convic to: acried out using the wide-Hield images on the tw routinely detect EGG to r = 25.0. Our survey sets to the sensitivity of the survey, we are able to sit principal lines detected in SIACT are hel indehind detail the properties of the survey as well as pre 535 ELG candidates in an area of 1.50 dog ⁺ (surf).	IVN 3.5 m telescope. Because of the s invations are made using three custom nettaneously detect sources via a numi rup to 0.144(), [D III]35007 (redshifts up tech initial results obtained by analytic rice density of 355 ELGs deg ²). Follow-	uperior depth and ex narrowband filters o ber of different emis- to 0.500], and [O II] it ig our three pilot stu- up spectra for a subs	cellent image quality afforded by WiYN, v entered on 6500 A, 6550 A, and 7460 A. Dr ion lines over a wide range of redshifts. If 727 (redshifts up to 1.015). In this paper, v dy fields. These fields have yielded a total

4. XCDS Version

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HD 142527 B" (<a href="https://doi.org/10.3847/1538-3881/ac73f4" title="2022, AJ, 164, 1">2022, AJ, 164, 1</a>)</title>
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5. HTML Rendering

2023AJ....166..135B AJ, volume 166, article 135, pages 1-4 published on the 01th of September 2023 by The American Astronomical Society, doi:10.3847/1538-3881/aced06

ERRATUM for doi:10.3847/1538-3881/ac73f4 3mfab v@

Erratum: "Improved Orbital Constraints and Hα Photometric Monitoring of the Directly Imaged Protoplanet Analog HD 142527 B" (2022, AJ, 164, 1)

Balmer William O. 1 2 0, Follette Katherine B. 3 0, Close Laird M. 4 0, Males Jared R. 4 0, De Rosa Robert J. 3 0 , Adams Redai Jéa I. 6 0, Watson Alex 3, Weinberger Alycia J. 7 0, Morzinski Katie M. 4 0, Morales Julio 8 0, ► ... (2 more authors)

► Affiliations...

Keywords:

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 1. Updated Photometry

 2. Mutual Inclination Update

 Acknowledgments

 References

This erratum corrects two errors in the published article [Balmer et al. 2022, First, we implement two corrections to our published photometry: we traction pipeline for by photometric extraction and calculation of the fir haminosity of the companion, hold on which now follow the proceedure described in detail in Follette et al. (2022). The updated photometric calibration and BKA model fitting results in astrometry that is identical to reported in Balmer et al. (2022) to within uncertainties, so we do not proget the orbital analysis here. We however make a small correction that arose from a numerical error in the calculation of the munual inclination angle between the biary orbit and the line are aloue effsk planes.

Neither correction results in a change to the main conclusion of the paper, namely the orbital solution for HD 142527 B. The evidence for a near perpendicular orientation for the binary relative to both disk planes remains strong. The updated photometry, which manifests primarily in an increase in photometric uncertainty does, however, cast into doubt the tentative claim made in the original article of photometric (and therefore accretion realy variability in the excess He mission of HD 142527 B.

1. Updated Photometry

In Balmer et al. (2022), we normalized the input point-spread function (PSF) when conducting PSF forward modeling, but did not also normalize the data cube that would be startight subtracted. This resulted in a PSF model without a fixed contrast telative to the star, a result of temporal variability in the PSF. In Follette et al. (2022) this normalization was implemented, and here we present corrected photometry for HD 14227 B.

Additionally, in the original article we used the formula,

□ 6. Technical requirements 1/2

• Currently:

• the BCS daily fetches from an FTP server an archive containing all articles updated the day before for all journals

• Problems:

- Issues sometimes incomplete (though they are marked as complete on the website)
- No easy access to old articles
 - Need to keep all updated articles on our server
 - Ask IOP for missing/old articles

Ideal solution:

• a Web service/API to ask for specific articles, issues and volumes (as the STACKS did in the past)

□ 6. Technical requirements 2/2

• Additional problem:

• *Errors in author names* (e.g. switched firstname and lastname, part of the firstname inside the lastname)

author	Ι	JANARDHAN	Ρ.
lastname	Ι	Ρ.	
firstname	Ι	Janardhan	
orcid	Ι	0000-0003-	-2504-2576

 Wrong ORCID (i.e. not the one of the author) orcid | 0000-0002-0786-7307 authors | BARRO G.,KELLY P.L.,LU L.-Z., | SCHWARZ G.J.,WILLIAMS T.G.,WU H.

7. Good points

- Usage of a standard XML schema (i.e. JATS) to format articles in XML
- Good communication with the IOP team
 - quick answers in case of problem (e.g. missing articles)
 - always get notifications in case of change on their side