CDS Gaia tutorial

Tutorial "Access to Gaia in CDS services" - Gaia data workshop @ Heidelberg

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For easy access to URLs and easy copy/paste, an online version of the instructions is available at <u>http://cds.unistra.fr/~boch/gaia-workshop-nov2016/</u>.

A <u>PDF version</u> is also available.

Each section (numbered 1. to 7.) can be done independently. Sections 1 to 6 do not require any prerequisite ; section 7 requires some Python knowledge.

1. Explore Gaia DR1 in Aladin Lite: density map and HiPS

Open <u>http://cds.unistra.fr/Gaia/DR1/AL-visualisation.gml</u> in your browser. By default, it shows a density map of Gaia DR1 sources. You can zoom in, zoom out and pan around.

For now, switch off the layer Gaia DR1 sources

Notice the satellite scanning patterns near the galactic center.



Click on the icon to jump to a given target. Type in *Andromeda* and submit. Zoom in if needed. The Andromeda galaxy is clearly visible.

Repeat the same process for objects *M* 33 and *NGC* 55. You can switch the base image layer to DSS or 2MASS and see how these galaxies look like in optical or infrared.



Switch to the DSS image layer, and jump to LMC (Large Magellanic Cloud). Adjust zoom

to frame the whole galaxy.



Switch on the *Gaia DR1 sources* HiPS progressive catalogue. When the field of view is large, only the brightest sources are displayed.

Click on one of the source and see the associated measurements displayed in the right panel.

As you zoom in, fainter stars appear.

Click on another object and select *More details* to see the VizieR page with the whole set of measurements.

NB: the same steps can be performed in Aladin Desktop. Type

<u>http://cds.unistra.fr/Gaia/DR1/hips/GaiaSourceDR1-density-map-8192/</u> in Aladin command bar and submit to load the Gaia density map. HiPS progressive catalogues can be loaded from the HiPS form (menu Fille —> Open, tab HiPS, then open node Catalog)

2. Query Gaia tables in VizieR

2.1 Query table I/337/gaia

Open VizieR homepage <u>http://vizier.u-strasbg.fr/viz-bin/VizieR</u> in your browser, type *Gaia* in the first text field and submit

Click on I/337/gaia

Type Cen A as the target name and 20 arcmin for the search radius and submit

Simple Target List Of Targets			
Target Name (resolved by Sesame) or Posi	tion:		Target dimension:
Clear M 44	J2000	\$ 20	arcmin 🖨 Submit
		(🖸 Radius 🔵 Box size

By default, VizieR outputs the result an HTML table, limited to 50 rows. This limit, as well as the output format, can be modified from the left sidebar shown below:

Preferences	
max: 50	٥
HTML Table	0
 All columns 	
Compute	
	Submit

Change the maximum to 9999 and resubmit to get the whole list of rows in the requested cone.

will show the sky positions of the rows in Aladin Lite.



plot the output Clicking on

attributes.

DF KOR



In our resut, only default columns are visible. Click on Modify guery, select a few more columns and Submit. Check that the selected columns are visible in the output.

We are now going to add a constraint on our positional query, in order to keep only sources with parallax measurement greater than 5 mas. Click again on *Modify query* and add the constraint:

	0	Plx	>5	mas

Submit. You should retrieve 14 sources matching the constraint.

This result can be easily transmitted to Virtual Observatory tools:

• Fire up a SAMP-compatible application (either Topcat or Aladin).



- Click on the antennae •

at the top right corner of the page

accept connection •



click on Broadcast •

2.2 Query table I/337/cepheid

Go back to http://vizier.u-strasbg.fr/viz-bin/VizieR, search for Gaia and select the Cepheids table 1/337/cepheid

Submit, change the max. number of rows to Unlimited as to retrieve all the 599 rows.

Click twice on the arrows above P1 label to sort the rows according to the period, in ascending order.



The first row is now the star with the shortest period. Clicking on the corresponding fov link will show the different measurements for this star.

Click on LC to display the light curve of this star.



We will now fold according to the cepheid period :

- click on the X axis label
 ObsTime [d] from T0=J2010.0
- click on and enter period (0.81104349), as read in the table





3. Retrieve Gaia measurements for stars in globular clusters, selected from **Simbad**

Open http://simbad.u-strasbg.fr/ and click on Queries by criteria

Type the following search expression
otype = '*inCl' & Bmag>11 & Bmag<12 & nbref>5

This query will search for stars in cluster having magnitude B between 11 and 12 and cited in more than 5 papers

Submit query. 1638 objects match the criteria

Number of objects: 1638

Go back to the query page, select Display and resubmit the query.

Return :			
odisplay	maximum 10000 🗘 objects		
or get references from the selected objects			

Go to the bottom of the result page, and store result in the CDS Portal

Store this result in Votable, in Ascii, or in the CDS portal

On the following page, change the filename and click on Save

Filename:	simbad-stars-in-cluster
Comment:	

We can now use this result as input for the cross-match service.

Go to http://cdsxmatch.u-strasbg.fr/

Save

At the top of the page, click on *Table management* and on *Add metadata* for the table we previously saved.

Change unit to sexa (sexagesimal) for both RA and DEC and click Update.

Manage x-match metadata of table simbad-stars-cluster	8
General metadata	
RA RA sexa Dec DEC deg For each kind of metadata, select the matching column from the table and its unit. Error metadata	
Error type No error	
No metadata to fill (1) For each kind of metadata, select the matching column from the table and its unit.	
Update Cance	1

tab

Go back to the

Choose simbad-stars-cluster as the 1st table and GAIA DR1 TGAS as the 2nd one.

Choose tables to cross-match	
simbad-stars-cluster VizieR SIMBAD My store	GAIA DR1 TGAS VizieR SIMBAD My store
simbad-stars-cluster 1,638 rows	Gaia DR1 (Gaia Collaboration, 2016) 2,057,050 rows

Open the *Options* panel, change *Radius* to *1 arcsec* and submit by clicking on *Begin the X-Match*

You can follow computation progress in the bottom table. Once the job is completed, click on the info icon to get a summary of the result and click on the histogram icon to see the distribution of distances for found associations.



Download the result as a VOTable, and load it in Topcat (*File* -> *Load table*) Create an histogram (*Graphics* -> *Histogram Plot*) and visualize the distribution of the *parallax* in our sample.



4. Gaia data in Aladin Desktop

In this part, we will learn how to access to and visualize Gaia data from Aladin Desktop.

Launch Aladin with the following command:

java -Xmx1024M -jar Aladin.jar -beta

4.1 Visualize proper motion of Gaia stars in Aladin Desktop

In the location bar at the top of the main Window, type *Praesepe* and submit. Zoom out to see the whole open cluster

Open the server selector: menu $File \rightarrow Open$



Gaia DR1 can be accessed from a dedicated tab

Click on the Gaia tab, change radius to 2°, select table I/337/tgasptyc and Submit.

From menu Catalog, select Create a scatter plot and plot pmRA vs pmDE. Stars which are member of Praesepe are clustered around (pmRA=-35mas/yr, pmDE=-12mas/yr)

Select these points in the scatter plot and notice they are automatically selected in the spatial view.



Go back to a single panel (menu View -> Panels -> 1 panel)

Create a filter by clicking on the icon filter and select the predefined filter *Draw proper motions of stars.*

To make the arrows bigger, switch to *Advanced mode* edit the filter as below and *Apply*: {

```
draw white plus
draw pm(10*$[pos.pm;pos.eq.ra], 10*$[pos.pm;pos.eq.dec])
```

Size of arrows can also be controlled with a slider available in the stack after selecting the Gaia catalogue plane:

Æ	🛋 Filter0
	💁 Gaia
v /@	🖲 DSS colored 👘 O
J2000	+
size	+
opac.	
zoom	+

Select a few sources by drawing a rectangle. Associated measurements appear at the bottom of the window:

2	📌 _RAJ2000	_DEJ2000	_V	TYC	HIP	Source	RA_ICRS	e_RA	DE_ICRS	e_DE	Plx	e_Plx	pmRA	e_pmRA	pmDE ¢
	129.6049097	21.2526305	VizieR	1398-1997-1		664744732785677056	129.6048809442	0.354	21.2526445233	0.12	1.68	0.89	-6.437	3.041	3.368
	130.7211003	20.8192389	VizieR	1399-1306-1	42766	664845303740260992	130.7209334337	0.273	20.8191802958	0.164	5.31	0.31	-37.438	0.081	-14.054
	130.7217203	20.8974635	VizieR	1399-1155-1		664849152030957056	130.7216811425	0.302	20.8974006679	0.149	2.35	0.84	-8.778	2.986	-15.071
	130.9713256	21.1779188	VizieR	1399-1574-1		664861143579642752	130.971275802	0.211	21.1779127841	0.132	1.9	0.41	-11.153	1.292	-1.433
	130.6370591	21.1788744	VizieR	1398-1797-1		664869905312917248	130.6370942179	0.384	21.1788399182	0.174	2.22	0.91	7.856	3.122	-8.279
	131.1743821	21.1033794	VizieR	1399-1178-1		664879251161368832	131.1743508377	0.38	21.1033819252	0.574	2.34	0.67	-6.989	1.312	0.612
	130.3687735	21.2281063	VizieR	1398-2317-1		664919589494596480	130.3687460395	0.412	21.2281753854	0.177	3.8	0.98	-6.133	3.325	16.587

As the catalogue contains positions, proper motions along with the epoch of observation, Aladin can compute the positions at a different epoch:

• select the Gaia plane and click on the icon to open the *Properties window*

:=

• move the slider to visualize stars position at the corresponding epoch

Epoch:	J2122	Img epoch	Reset
	17001800190@200@100	22002300	

4.2. Retrieve Gaia sources within HST coverage

Goal: In this subsection, we will learn to use MOCs in order to retrieve Gaia data which are in the common coverage between HST and SDSS observations.

Open the server selector (*File* \rightarrow *Open*) and click on the HiPS tab. Expand the nodes *Image* \rightarrow *Optical* \rightarrow *SDSS* and select *SDSS9 colored*. Expand also nodes *Image* \rightarrow *Optical* \rightarrow *SDST* and select *HST-R* (Canadian Astronomy Data Centre)

Submit.

The 2 corresponding HiPS are loaded. For each HiPS plane:

- Right-click on its label, select Properties
- Click on Coverage to load the corresponding MOC

Coverage		
Space	35.62 % of sky	Coverage

Open menu *Coverage* —> *Logical operations*, select the 2 MOCs and compute intersection.

••	MOC operations
	Specify one or two MOC planes, choose a MOC operation and press the CREATE button to generate the resulting MOC.
Plane	HST-R MOC - "01 30 00.00 +24 37 27.5"
Plane	SDSS9 u MOC - "01 30 00.00 +24 37 27.5"
Plane	none 📀
Plane	none 😒
Plane	none 📀
Plane	none
OUr	ion 💿 Intersection 🔿 Subtraction 🔿 Difference 🔿 Complement
	CREATE Reset Close ?

The resulting MOC plane is the common coverage between SDSS9 and HST-R. VizieR tables can be queried to retrieve only sources within a given MOC:

- open the server selector
- select the *MOC* tab, in the right hand panel
- select the MOC plane we just created
- enter following VizieR table ID: GAIA-DR1
- change maximum number of rows to unlimited
- submit

Zoom in and pan around to verify that returned rows fall into the MOC coverage.

5. Access to CDS cross-match service from Topcat : find Gaia and 2MASS couterparts for a list of sources and draw a color-color diagram

Download on your machine the table located at <u>http://cds.unistra.fr/~boch/gaia-workshop-nov2016/data/allwise-SMC-sample.vot</u> It is a sample of 313k AllWISE sources located in the Small Magellanic Cloud.

Launch topcat: java -Xmx1024M -jar topcat-full.jar

Load the table (*File* -> *Load Table*)

Open the CDS X-Match window from VO —> CDS Upload X-Match Select Gaia DR1 in the list of VizieR tables Select allwise-SMC-sample.vot as the input table. Set the match radius to 1 arcsec and submit.

	CDS Upload X-Match						
-	K						
Remote Table-							
VizieR Table ID	/Alias: GAIA DR1						
Name: I/337/gaia							
Alias: G	Alias: GAIA DR1						
Description: GaiaSource data ({\bf Download } Gaia Source							
Row Count: 1 142 679 769							
Coverage: 0.	9999797 (order 6) 🛛 👥 💿						
Local Table							
Input Table 1: allwise-SMC-sample.vot ᅌ							
RA column:	AJ2000 🔻 degrees ᅌ (J2000)						
Dec column: DEJ2000 🔻 degrees 🗘 (J2000)							
r Match Parameters							
Radius: 1.0	arcsec 🗘						
Find mode:	Best 🗘						
Rename columns: Duplicates ᅌ Suffix: _x							
Block size: 50000 💌							
	Co Stop						

Once the cross-match is done, open the result table and notice that each row contains the fields of the original file plus the fields coming from the Gaia data.

Click on icon and create the histogram for column *angDist* (separation in arcsec between AllWISE source and Gaia source) as to assess the quality of the match.



In order to retrieve J, H, K magnitudes, we will now search for counterparts in 2MASS. Open the *CDS X-Match* window, select 2MASS in the list of VIzieR tables and *1xGAIA DR1* as input table.

CDS Upload X-Match					
	×				
Remote Table					
VizieR Table I	D/Alias: 2MASS				
Name:	II/246/out •				
Alias:	2MASS				
Description:	The Point Source catalogue of 470,992,970 sc				
Row Count:	470 992 970				
Coverage:	0.9999797 (order 6) 🛛 🔵 🔍				
Local Table					
Input Table	2: 1xGAIA DR1				
RA column:	RAJ2000 🔻 degrees 🗘 (J2000)				
Dec column:	DEJ2000 v degrees 🗘 (J2000)				
Match Parame	ters —				
Radius: 1.0	arcsec ᅌ				
Find mode:	Best 🗘				
Rename colur	nns: Duplicates ᅌ Suffix: _x				
Block size: 50000 🔻 🕘					
Co Stop					

We should retrieve 185,417 rows.

Create a scatter plot for this result (*Graphics* —> *Plane Plot*) and plot **X:** *phot_g_mean_mag* - *Jmag* vs **Y:** *Jmag* - *W1mag*



6. More complex queries with TAP VizieR

6.1 Select high proper motions stars

Goal: retrieve 10 TGAS stars with higher proper motion near the galactic center

Go to TAP VizieR web interface: <u>http://tapvizier.u-strasbg.fr/adql/</u>

Search tables for the Gaia keyword.

Select I/337/tgas and click on Construct your query at the top right

Click on *Sky area* and enter *Galactic center* as cone center Update radius to 5 *deg, max records* to 10 and click on *Update query* Click on *Quickview* to get a preview of the result

Edit the query to keep only source_id, ra, dec and compute total proper motion: SELECT TOP 10 source_id, ra, dec, sqrt(pmra*pmra+pmdec*pmdec) as pm FROM "I/337/tgas" WHERE 1=CONTAINS(POINT('ICRS', "I/337/tgas".ra, "I/337/tgas".dec), CIRCLE('ICRS', 266.416833, -29.007806, 5.))

Click on Quickview to get a preview

We still need to sort the result according to the proper motion value, using the ORDER BY clause.

```
The final query should look like this:

SELECT TOP 10 ra, dec, sqrt(pmra*pmra+pmdec*pmdec) as pm

FROM "I/337/tgas"

WHERE 1=CONTAINS(POINT('ICRS',"I/337/tgas".ra,"I/337/tgas".dec),

CIRCLE('ICRS', 266.416833, -29.007806, 5.))

ORDER BY pm DESC
```

Click on Run to launch the query and retrieve the result in the requested format.

6.2 Recreate HR diagram from Gaia DR1 paper

Goal: we will recreate figure 3c of the Gaia Data release 1 paper (<u>https://arxiv.org/pdf/1609.04172v1.pdf</u>)

Launch Topcat: java -Xmx1024M -jar topcat-full.jar

Open the TAP query window (menu VO -> Table Acces Protocol (TAP) Query)

Select TAPVizieR (31971) - ivo://cds.vizier/tap

and click on

Use Service

In the appendix of https://arxiv.org/pdf/1609.04172v1.pdf , table B.1 provides with the ADQL query used to create the HR diagram:

```
SELECT gaia.source_id, gaia.hip,
gaia.phot_g_mean_mag+5*log10(gaia.parallax)-10 as g_mag_abs,
hip.b_v
FROM gaiadr1.tgas_source as gaia
inner join public.hipparcos_newreduction as hip
on gaia.hip = hip.hip
WHERE gaia.parallax/gaia.parallax_error >= 5 and hip.e_b_v > 0.0
and hip.e_b_v <= 0.05 and</pre>
```

2.5/log(10)*gaia.phot_g_mean_flux_error/gaia.phot_g_mean_flux <= 0.05

This query is meant to be executed on GACS archive at ESA. A few changes must be made in order to make it work with TAP VizieR:

- In TAP VizieR, gaiadr1.tgas_source is named ... and public.hipparcos_newreduction is named ...
- columns b_v and e_b_v respectively are named B-V and e_B-V

As the dash is a special character in ADQL, the column name must be put between double quotes: "B-V" and " e_B-V "

• column hip is named HIP

We will also add position fields ra and dec from TGAS

```
The updated ADQL query to be typed in Topcat is:
SELECT gaia.ra, gaia.dec, gaia.source_id,
gaia.hip, gaia.phot_g_mean_mag+5*log10(gaia.parallax)-10 as
g_mag_abs,
hip. "B-V"
FROM "I/337/tgas" as gaia
inner join "I/311/hip2" as hip
on gaia.hip= hip.HIP
where gaia.parallax/gaia.parallax_error >= 5 and hip. "e_B-V" > 0.0
and hip. "e_B-V" <= 0.05 and
2.5/log(10)*gaia.phot_g_mean_flux_error/gaia.phot_g_mean_flux <=
0.05
```

Copy/paste this query in the ADQL text panel, at the bottom of the window and *Run Query:*

ADQL Text	
SELECT gaia.source_id, gaia.hip, gaia hip."B-W" FROM "I/337/tgas" as gaia inner join "I/31/hip2" as hip on gaia.hip= hip.HIP where gaia.parallax/gaia.parallax_err 2.5/log(10)+gaia.phot_g_mean_flux_err	a.phot_g_mean_mag+5+log10(gaia.parallax)-10 as g_mag_abs, ror >= 5 and hip."e_B-V" > 0.0 and hip."e_B-V" <= 0.05 and ror/gaia.phot_g_mean_flux <= 0.05
Examples • •	Info
Examples ()	Run Query

We should retrieve 74,817 sources.

Create a scatter plot (*Graphics* -> *Plane Plot*) and select *B-V* for **X** axis and *g_mag_abs* for **Y** axis.





6.3 Retrieve spectral types from Simbad

We will now try to retrieve the spectral types of our sources, by querying Simbad through the CDS cross-match service.

Click on , select *simbad* in remote table list. Select *TAP_1_I_337_tgas,I_311_hip2* as input table and launch cross-match at 1 arcsec.

On the result table, we will create some subsets according to the spectral type of the star.

Click on and define a new subet named *O stars* matching the following expression: sp_type.startsWith("0")



Repeat the same steps to create subsets for spectral types **B**, **A**, **F**, **G**, **K** and **M**.

Once the 6 spectral types have been created, plot the HR diagram following steps of section 6.2

Click on the Subsets tab, unselect All and select all 6 subsets:

			_	\$		AII
			_	\$		D :
			_	ŧ	 Image: A start of the start of	3
			_	ŧ	 Image: A start of the start of	A
			_	\$	✓	F
Position	Subsets	Form	_	\$	 Image: A start of the start of	G

You should end up with the following plot:



7. Accessing CDS services with Python

In this part, we will redo part of the previous sections, but in a programmatical manner, using Python.

A static version of the notebook is available at <u>https://github.com/tboch/VO-access-GaiaDR1/blob/master/notebooks/gaia-cds-services.ipynb</u>, providing examples on how to access Gaia data from CDS services.

If you want to modify the notebook and adapt queries to your own use cases, you can either:

- download the iPython notebook locally to run on your own laptop (requirements: *Python3* and libraries *astropy, astroquery, numpy, mocpy*).
- run it from mybinder. No need to install anything and you can still interact/update the notebook.